

¹ **On the origins of temporal power-law behavior in the**
² **global atmospheric circulation**

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4 Climate variations on timescales longer than a year are often character-
5 ized by temporal scaling (“power-law”) behavior for which spectral power
6 builds up at low frequencies in contrast to red-noise behavior for which spec-
7 tral power saturates at low frequencies. Checks on the ability of climate pre-
8 diction models to simulate temporal scaling behavior represent stringent per-
9 formance tests on the models. We here estimate temporal power-law expo-
10 nents (“Hurst exponents”) for the global atmospheric circulation of the strato-
11 sphere and troposphere during the 20th century. We show that current-generation
12 climate models generally simulate the spatial distribution of the Hurst ex-
13 ponents well. We then use simulations with different climate forcings to ex-
14 plain the Hurst exponent distribution and to account for discrepancies in scal-
15 ing behavior between different observational products. We conclude that char-
16 acterization of temporal power-law behavior provides a valuable tool for cross-
17 validating low-frequency variability in various datasets, for elucidating the
18 physical mechanisms underlying this variability, and for statistical testing
19 of trends and periodicities in climate time series.

1. Introduction

Climate variability on interannual to multi-decadal time scales involves a mix of externally and internally generated variability [Wigley and Raper, 1990]. The classical two-parameter model of such variability is Hasselmann [1975] autoregressive model of the first order (AR1). It corresponds to a class of physical models in which stochastic (weather-noise) atmospheric variability drives slower components of the climate system such as the ocean. An alternative two-parameter model of the temporal power spectrum is the power-law model

$$S_{PL}(\lambda) = b|\lambda|^{1-2H}, \quad 0 < \lambda_l \leq |\lambda| \leq \lambda_h \leq 1/2, \quad (1)$$

where λ is the frequency, b represents the overall spectral power, the “Hurst exponent” H is related to the spectral slope s by $H = (1-s)/2$, and λ_l and λ_h are low and high frequency cutoffs used in model fitting. Unlike the Hasselmann model, the power-law model, which indicates temporal scaling behavior rather than dependence on any particular timescale, has no simple established physical interpretation.

Recent research has pointed out potential limitations of the AR1 model [e.g. Hall and Manabe, 1997] and has shown that power-law scaling behavior arises in surface air temperature [Pelletier, 2002; Blender and Fraedrich, 2003; Huybers and Curry, 2006], the atmospheric circulation [Tsonis et al., 1999; Vyushin and Kushner, 2009], etc. The current instrumental record is too short to statistically claim the superiority of the one model over the other on timescales shorter than a century, but there are locations where power-law seems to fit the observations better than AR1 [Percival et al., 2001; Vyushin et al., 2007; Vyushin and Kushner, 2009]. We therefore do not claim that power-law behavior

33 is universal on all timescales, and instead use $S_{PL}(\lambda)$ to provide a sense of how quickly
34 power builds towards lower frequencies. Regions where $\hat{H} = 0.5$ (the flat spectrum limit)
35 might be well described by either model, while regions where \hat{H} is closer to 1 (the $1/f$
36 limit) are candidates for true power-law behavior.

37 We here report on how climate prediction models can be used to simulate and explain
38 the observed spatial distribution of the Hurst exponent estimate \hat{H} for the atmospheric
39 general circulation. To do so, we compare \hat{H} from observationally based reanalysis prod-
40 ucts to comprehensive climate simulations, and then use more specialized simulations to
41 explain specific features of the \hat{H} field. Previous model-observation comparisons have
42 concluded that the ability of climate models to simulate the observed scaling is mixed
43 [*Govindan et al.*, 2002; *Blender and Fraedrich*, 2003; *Vyushin et al.*, 2004], but this work
44 has generally been restricted to surface air temperature and has proven to be method and
45 model dependent. We here carry out physically motivated analyses and provide cross-
46 validation checks that are independent of the Hurst exponent estimation technique. In
47 separate work, we have verified that alternative Hurst exponent methods provide similar
48 results [*Vyushin and Kushner*, 2009].

2. Data and Methods

49 To estimate H for $S_{PL}(\lambda)$ in (1), we use detrended fluctuation analysis of the third
50 order [DFA3, *Kantelhardt et al.*, 2001]. See the supplementary information for DFA3
51 details and a comparison of its results to another method. We estimate H for the zonal-
52 mean air temperature data in the range of 18 months to 45 years for the NCEP/NCAR
53 and ERA40 reanalyses for the period 09.1957-08.2002. We compare these estimates to

54 several model simulations: simulations of the GFDL AM2.1-LM2.1 atmospheric general
55 circulation model [*The GFDL Global Atmospheric Model Development Team, 2004*], for
56 which sea surface temperatures (SSTs) are prescribed, and 17 coupled model runs of the
57 20th century from the CMIP3 archive. All model simulations were taken for the period
58 01.1955-12.1999. Additional model details are provided in the supplementary informa-
59 tion. Because the quasi-biennial oscillation (QBO) is not captured by any of the models
60 considered, we filter the QBO signal from the reanalyses temperature in addition to the
61 seasonal cycle [*Vyushin et al., 2007*].

3. Results and Discussion

62 Fig. 1 plots estimates of H for the reanalysis products and several climate simulations.
63 The \hat{H} distribution displays a characteristic shape that we have verified is robust to
64 different methods of H estimation [*Vyushin and Kushner, 2009*]. Both the NCEP/NCAR
65 and ERA40 reanalyses (Figs. 1a and b) show maxima in \hat{H} in the tropical to low-
66 extratropical troposphere and in the tropical to subtropical stratosphere and a minimum
67 in the Northern Hemisphere polar stratosphere. But there are differences between the
68 reanalysis products; for example, ERA40 has separate local maxima in \hat{H} in the lower
69 and upper troposphere at 60°S that will be discussed later in relation to Fig. 3. We will
70 also show that even where the distributions appear to agree, they might do so for different
71 reasons.

72 Fig. 1c plots the \hat{H} distribution for a simulation of AM2.1 forced by historical SSTs, an-
73 thropogenic greenhouse gases and aerosols, ozone changes, solar flux, and volcanic aerosols
74 (hereafter the “HistSST+AllForc” simulation). The main features of the \hat{H} distribution

75 of this simulation are similar to that displayed in the observationally based Figs. 1a-b,
76 including the falloff of \hat{H} as we move from the equator to the poles and separate maxima
77 in the lower stratosphere and the troposphere. Therefore given historical SSTs and the
78 other principal external forcings GFDL AM2.1 is able to reproduce the continuum of zonal
79 mean temperature variability represented by the Hurst exponent.

80 We use three additional simulations of AM2.1 to explain the \hat{H} distribution: 1) “Climo”,
81 a simulation in which the climate forcings including SSTs are not allowed to vary from
82 year to year. The “Climo” simulation has \hat{H} values fairly close to 0.5 everywhere, with
83 a range of between 0.4 and 0.6 (not shown). This supports Hasselmann’s assumption of
84 a flattening out of the spectrum at low frequencies, demonstrating an absence of long-
85 term memory in the atmosphere in the absence of coupling to the ocean; 2) “HistSST”,
86 a simulation that is forced with historical SSTs but that keeps all other climate forcings
87 fixed. The “HistSST” simulation (Fig. 1d) gives rise to a tropospheric pattern of \hat{H}
88 that is similar to that in Figs. 1a-c. This is consistent with our classical understanding
89 that the tropospheric circulation and thermal structure are largely determined once the
90 SSTs are prescribed on timescales longer than the atmospheric adjustment timescale of a
91 few months; 3) “Vol”, a simulation that is forced with the “Climo” SSTs but that uses
92 historical volcanic forcing, while keeping all other forcings fixed. The “Vol” simulation
93 (Fig. 1e) gives rise to a stratospheric pattern of \hat{H} that is similar to that in Figs. 1a-c.
94 Thus, the simulations show that the observed \hat{H} distribution is mainly determined by
95 temporal variability of the SSTs in the troposphere and by volcanic forcing in the lower
96 stratosphere.

97 We briefly demonstrate that current generation climate models can capture the \hat{H} dis-
98 tribution in a less constrained forcing framework. The \hat{H} distribution averaged over the
99 CMIP3 coupled ocean-atmosphere models is shown in Fig. 1f; it displays a similar struc-
100 ture to Figs. 1a-c but has a narrower meridional extent and a weaker volcanic signature
101 in the lower stratosphere. The simple explanation for the latter is that only 9 of the 17
102 models considered included realistic volcanic forcings.

103 We propose that the relatively steep spectral slopes represented by the \hat{H} maximum
104 centered in the tropical troposphere are generated by tropical SST variability. Our test
105 of this idea reveals a significant discrepancy between the two reanalysis products. To
106 test the idea, we create time series of tropical mean SST in the latitude band 20⁰S-
107 20⁰N (“TropSST”). We then filter the TropSST signal from the temperature time se-
108 ries and estimate H of the result for the NCEP/NCAR and ERA40 reanalyses and for
109 the HistSST+AllForc simulations. Fig. 2 isolates the part of the \hat{H} distribution re-
110 lated to tropical SSTs by showing the original \hat{H} minus the TropSST-filtered \hat{H} . In the
111 NCEP/NCAR reanalysis (Fig. 2a) and in the simulation (Fig. 2c), there is a vertically
112 coherent part of the \hat{H} distribution throughout the tropical and low extratropical tropo-
113 sphere that is related to the TropSST signal, as indicated by the positive values. The
114 TropSST \hat{H} signature in the ERA40 reanalysis (Fig. 2b) is qualitatively different, being
115 vertically incoherent and of mixed sign.

116 In Fig. 2, the NCEP/NCAR reanalysis and the climate model simulation appear to
117 agree with our hypothesis of tropical SST control, while the ERA40 appears to disagree
118 with it. To understand these inconsistent results we display the residuals of the tropical

119 upper tropospheric temperatures after TropSST filtering has been applied, for the reanal-
120 ysis products and for the HistSST+AllForc and HistSST simulations (Fig. 3a). A one year
121 running average has also been applied. The ERA40 residuals (shown in red) show much
122 more decadal variance than the NCEP/NCAR residuals and the simulations' residuals.
123 Significant fluctuations for the ERA40 include particularly high values during 1975-1983,
124 which are probably related to problems with transition from VTPR to TOVS satellite data
125 [*Simmons et al.*, 2004; *Uppala and Coauthors*, 2005], and low values after 1992. Similar
126 issues also explain the lower and upper tropospheric \hat{H} maxima at 60°S that are seen in
127 the ERA40 reanalysis (Fig. 1b) but not seen in the NCEP/NCAR reanalysis (Fig. 1a) or
128 in the HistSST+AllForc simulation (Fig. 1c). Figs. 3b and c plot temperature anomalies
129 (without TropSST filtering) from the same four data sets at these locations. There is an
130 obvious jump (negative at 925hPa and positive at 300hPa) in the ERA40 temperature
131 presumably related to problems with assimilation of the VTPR data from 1973 to 1978
132 [*Bengtsson et al.*, 2004; *Simmons et al.*, 2004]. Another striking difference between the
133 models and reanalyses are the strong positive trends at 300hPa. These trends seem to
134 be spurious and stem from the reanalysis models cold biases combined with a gradual in-
135 crease in the amount of observations in the Southern Hemisphere [*Bengtsson et al.*, 2004;
136 *Simmons et al.*, 2004]. Discrepancies in the Southern Hemisphere polar stratosphere have
137 been discussed elsewhere [*Vyushin and Kushner*, 2009]. Therefore data inhomogeneity
138 issues in the ERA40 affect and are revealed by our H analysis.

4. Conclusions

139 To conclude, we find that zonal-mean air temperature on interannual to multi-decadal
140 timescales has a steep spectrum that might be modelled by power-law behavior in the
141 tropical to low-extratropical troposphere and the tropical to subtropical stratosphere.
142 Current generation climate models can capture these features and specialized forcing
143 simulations elucidate their dynamics. We propose that the tropospheric \hat{H} signatures
144 are linked to tropical SST variability and that the lower stratospheric \hat{H} signatures are
145 linked to volcanic forcing. The link to tropical SST variability is clear in only one of the
146 two observational products we use: the NCEP/NCAR reanalysis. The large \hat{H} values
147 in the tropical upper troposphere in the ERA40 reanalysis appears to arise from data
148 problems that mask the connection to tropical SSTs. The ERA40 H estimates also
149 exhibits tropospheric maxima at 60°S that appear related to other documented data
150 assimilation issues.

151 This analysis points to problems in naively interpreting the Hurst exponent distribution
152 as an indicator of long-term memory in climate and care needs to be taken to elucidate the
153 physical basis for a given \hat{H} feature. Data inhomogeneities affect many observational time
154 series and can equally give rise to power-law behavior [*Berton, 2004; Rust et al., 2008*].
155 Sometimes, such as at 60°S in the troposphere, it is immediately evident that there is a
156 discrepancy to explain, but at other times, such as in the tropical troposphere, the effort
157 still needs to be made to test the consistency of the power-law behavior under different
158 physical hypotheses. We have found that general circulation models provide a useful tool
159 for such testing.

160 The frequent presence of power-law behavior, whatever its cause, suggests that statis-
161 tical testing for significant trends and periodicities should use power-law noise models
162 [Smith, 1993; Vyushin et al., 2007] as well as AR1-models, particularly in the tropical up-
163 per troposphere and lower stratosphere where \hat{H} is large and trend evaluation has proven
164 difficult [e.g. Santer et al., 2005]. Power-law based confidence intervals are typically larger
165 because they assume more power at lower frequencies. For example, power-law based sig-
166 nificance testing has been applied to the problem of stratospheric ozone recovery in the
167 presence of significant stratospheric internal variability, and leads to a lengthening of the
168 projected time for the detection of ozone recovery [Vyushin et al., 2007].

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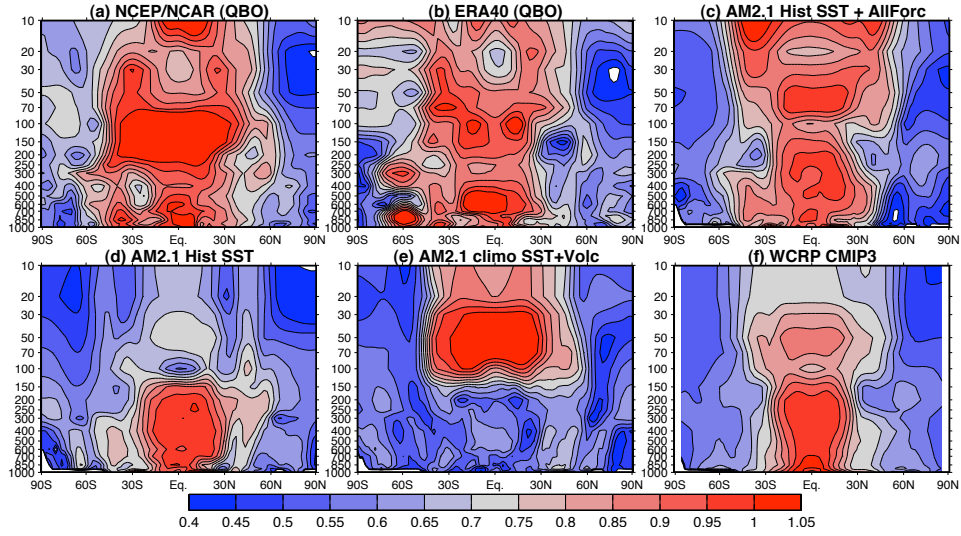


Figure 1. \hat{H} distribution for zonal-mean temperature for (a) the NCEP/NCAR reanalysis, (b) the ERA40 reanalysis, (c) the GFDL AM2.1 HistSST+AllForc simulation, (d) the GFDL AM2.1 HistSST simulation, (e) the GFDL AM2.1 Volc simulation, (f) the CMIP3 simulations. Panel (f) represents a multiple model average. As stated in the text, QBO filtering has been applied to the reanalysis temperatures in Figs. 1a-b.

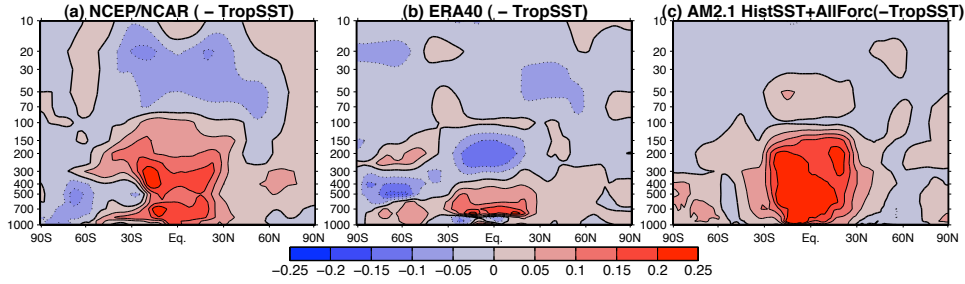


Figure 2. \hat{H} without TropSST filtering minus \hat{H} with TropSST filtering, which represents the signature of the tropical SSTs in the \hat{H} field: (a) NCEP/NCAR, (b) ERA40, (c) AM2.1 HistSST+AllForc. QBO filtering has been applied to ERA40 and NCEP/NCAR reanalyses.

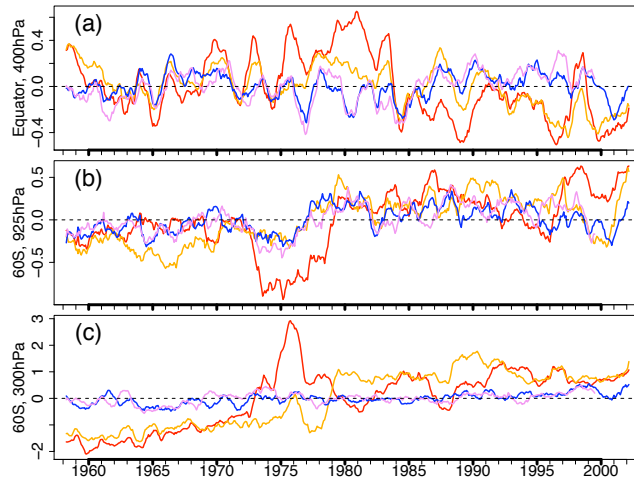


Figure 3. The one year running mean of zonally averaged air temperature residuals (a) at (Equator, 400 hPa) with TropSST filtering as described in the text; (b) at (60°S, 925 hPa), without TropSST filtering; (c) as in (b), at (60°S, 300 hPa). ERA40 time series are shown in red, NCEP/NCAR in orange, HistSST in blue, and HistSST+AllForc in violet.