# Solar Altitude and Azimuth Angle Calculation 

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## 1 Solar altitude angle calculation



Figure 1: Schematic of the Earth North Hemisphere showing spatial geometric relationships of interested locations: Solar altitude

Refer to figure 1, where, $O$ is the center of the Earth; $N$ is the north pole of the Earth; arch $N A Q$ and arch $N B T$ are two meridians of the Earth; arch $E Q T$ is the equator; $A$ is the location at the Earth surface where the solar altitude angle is to be calculated; $B$ is the location at the Earth surface from which the Sun looks like right at the zenith of the sky; $C$ is on line $O T$ with $B C \perp O T ; D$ is on line $O Q$ with $A D \perp O Q ; F$ is on line $A D$ with $B F \perp A D$. It is not difficult to tell that, $\measuredangle B O C$ is the solar decline angle, $\measuredangle A O D$ is the latitude of point $A, \measuredangle C O D$ is the hour angle before noon of point $Q$, and $\measuredangle A O B$ is the solar zenith angle of point $A$.

We know that the solar altitude angle is the complementary angle of the solar zenith angle. So the effort is to calculate the zenith angle $\measuredangle A O B$ at point $A$. This angle can be calculated by using the Law of Cosines:

$$
\begin{equation*}
c^{2}=a^{2}+b^{2}-2 a b \cos (\measuredangle C) \tag{1}
\end{equation*}
$$

where, $a, b$, and $c$ are side lengths of the triangle; and $\measuredangle C$ denotes the angle by side $a$ and $b$.

The Law of Cosines allows to calculate any of the three angle of a triangle if the lengths of the triangle's three sides are all known. For the triangle $\triangle A O B$ in figure (1), by setting the radius of the sphere to be unit, we have the length of 1 for both sides $A O$ and $B O$. The length of side $A B$ is to be calculated. We can tell the trapezoid $A B C D$ consists of a rectangle $B C D E$ and right triangle $A B F$. Now if we know the lengths of the right angle sides $A F$ and $B F$ of $\triangle A B F$, the length of the hypotenuse $A B$ will be known by the Pythagorean theorem. Now we use the name of a side/line denotes its length as well, then it can be seen
the following relations hold,

$$
\begin{align*}
B F^{2}= & C D^{2}=O C^{2}+O D^{2}-2 \times O C \times O D \times \cos (\measuredangle C O D) \\
= & {[B O \times \cos (\measuredangle B O C)]^{2}+[A O \times \cos (\measuredangle A O D)]^{2} } \\
& -2 \times[B O \times \cos (\measuredangle B O C)] \times[A O \times \cos (\measuredangle A O D)] \times \cos (\measuredangle C O D) \\
= & {[1 \times \cos (\measuredangle B O C)]^{2}+[1 \times \cos (\measuredangle A O D)]^{2} } \\
& -2 \times[1 \times \cos (\measuredangle B O C)] \times[1 \times \cos (\measuredangle A O D)] \times \cos (\measuredangle C O D) \\
= & {[\cos (\measuredangle B O C)]^{2}+[\cos (\measuredangle A O D)]^{2} } \\
& -2 \times \cos (\measuredangle B O C) \times \cos (\measuredangle A O D) \times \cos (\measuredangle C O D)  \tag{2a}\\
A F^{2}= & {[A D-D F]^{2}=[A D-B C]^{2} } \\
= & {[A O \times \sin (\measuredangle A O D)-B O \times \sin (\measuredangle B O C)]^{2} } \\
= & {[1 \times \sin (\measuredangle A O D)-1 \times \sin (\measuredangle B O C)]^{2} } \\
= & {[\sin (\measuredangle A O D)-\sin (\measuredangle B O C)]^{2} } \tag{2b}
\end{align*}
$$

To simplify the notations, define $\alpha=\measuredangle B O C, \beta=\measuredangle A O D, \Omega=\measuredangle C O D$, and $\psi=\measuredangle A O B$. With the new symbols, equations (2b)and (2b) are represented as,

$$
\begin{align*}
& B F^{2}=\cos ^{2} \alpha+\cos ^{2} \beta-2 \cos \alpha \cos \beta \cos \Omega  \tag{3a}\\
& A F^{2}=[\sin \beta-\sin \alpha]^{2}=\sin ^{2} \beta+\sin ^{2} \alpha-2 \sin \beta \sin \alpha \tag{3b}
\end{align*}
$$

From equation (3a) and (3b), by Pythagorean theorem, the length of $A B$ can be calculated as below,

$$
\begin{align*}
A B^{2} & =B F^{2}+A F^{2} \\
& =\left[\cos ^{2} \alpha+\cos ^{2} \beta-2 \cos \alpha \cos \beta \cos \Omega\right]+\left[\sin ^{2} \beta+\sin ^{2} \alpha-2 \sin \beta \sin \alpha\right] \\
& =\sin ^{2} \alpha+\cos ^{2} \alpha+\sin ^{2} \beta+\cos ^{2} \beta-2 \cos \alpha \cos \beta \cos \Omega-2 \sin \alpha \sin \beta \\
& =1+1-2 \cos \alpha \cos \beta \cos \Omega-2 \sin \alpha \sin \beta \\
& =2-2 \cos \alpha \cos \beta \cos \Omega-2 \sin \alpha \sin \beta \tag{4}
\end{align*}
$$

Now for triangle $\triangle A O B$, using Law of Cosine, yield,

$$
\begin{align*}
\cos \psi & =\frac{A O^{2}+B O^{2}-A B^{2}}{2 \times A O \times B O}=\frac{1^{2}+1^{2}-A B^{2}}{2 \times 1 \times 1}=\frac{1}{2}\left[2-A B^{2}\right] \\
& =\frac{1}{2}[2-(2-2 \cos \alpha \cos \beta \cos \Omega-2 \sin \alpha \sin \beta)] \\
& =\frac{1}{2}[2 \cos \alpha \cos \beta \cos \Omega+2 \sin \alpha \sin \beta] \\
& =\sin \alpha \sin \beta+\cos \alpha \cos \beta \cos \Omega \tag{5}
\end{align*}
$$

The hour angle $\Omega$ of point $Q$ in the above equation decreases with the rotation of the Earth during the day time. Since $\Omega$ is a function of the planet angular velocity $\omega$ of the Earth and local time $t$ of point $Q$, in order to conveniently use $\omega t$ to represent the hour angle, the following formula is considered,

$$
\begin{equation*}
\cos \Omega=\cos (\pi-\measuredangle Q O E)=-\cos (\measuredangle Q O E)=-\cos \omega t \tag{6}
\end{equation*}
$$

Substitute equation (6) into (5), get,

$$
\begin{equation*}
\cos \psi=\sin \alpha \sin \beta-\cos \alpha \cos \beta \cos \omega t \tag{7}
\end{equation*}
$$

The solar altitude angle, $\Psi$, is the complementary angle of zenith angle $\psi$, which gives,

$$
\begin{equation*}
\cos \psi=\cos \left(\frac{\pi}{2}-\Psi\right)=\sin \Psi \tag{8}
\end{equation*}
$$

Substitute equation (8) into (7), we finally get the solar altitude angle formula for point $A$ as below,

$$
\begin{equation*}
\sin \Psi=\sin \alpha \sin \beta-\cos \alpha \cos \beta \cos \omega t \tag{9}
\end{equation*}
$$

Where,

| $\Psi$ | $=$ solar altitude angle | $(\mathrm{rad})$ |
| ---: | :--- | ---: |
| $\alpha$ | $=$ solar declination angle | $(\mathrm{rad})$ |
| $\beta$ | $=$ site latitude | $(\mathrm{rad})$ |
| $\omega=$ | earth angular velocity | $\left(\frac{\pi}{12} h^{-1}\right)$ |
| $t=$ | local time $[0,24)$ | $(h)$ |

## 2 Solar azimuth angle calculation



Figure 2: Schematic of the Earth North Hemisphere showing spatial geometric relationships of interested locations: Solar azimuth

Refer to figure 2, where, $O$ is the center of the Earth; $N$ is the north pole of the Earth; $\operatorname{arch} N A Q$ and arch $N B T$ are two meridians of the Earth; arch $E Q T$ is the equator; $A$ is the location at the Earth surface where the solar azimuth angle is to be calculated; $B$ is the location at the Earth surface where the Sun looks like right at the zenith of the sky; $F$ is the projection of point $B$ on plane $A O Q$, e.g., $B F \perp A O Q ; C$ is on line $O T$ with $B C \perp O T ; D$ is on line $O Q$ with $F D \perp O Q$; by $B F \perp A O Q, B F \perp O Q$; by $F D \perp O Q$ and $B F \perp O Q, O Q \perp B C D F$; by $O Q \perp B C D F, C D \perp O Q ; G$ is on line $A O$ with $B G \perp A O$; by $B F \perp A O Q, B F \perp F G$ and $B F \perp A O$; by $B G \perp A O$ and $B F \perp A O, A O \perp B F G$; by $A O \perp B F G, F G \perp A O$; by $B G \perp A O$ and $F G \perp A O$, the angle $\measuredangle B G F$ is the angle by planes $A O Q$ and $A O B$, which is exactly the solar azimuth angle from the due south; meanwhile, it is easy to see that $\measuredangle B O C$ is the solar decline angle, $\measuredangle A O D$ is the latitude of point $A, \measuredangle C O D$ is the hour angle before noon of point $Q$, and $\measuredangle A O B$ is the solar zenith angle of point $A$.

Our effort is to calculate $\measuredangle B G F$. By $B F \perp F G$, we know the $\triangle B F G$ is a right triangle. For the right triangle $\triangle B F G$, if any two of the three sides are known, then all the three internal angles will be known. It can be seen that, the hypotenuse side $B G$ is the right angle side of another right triangle $\triangle B O G$, with $\measuredangle B O G$ being exactly the zenith angle of point $A$. Since the zenith angle becomes a known after using formula (9), then the length of $B G$ is,

$$
\begin{equation*}
B G=B O \times \sin (\measuredangle B O G)=\sin \psi=\sin \left(\frac{\pi}{2}-\Psi\right)=\cos \Psi \tag{10}
\end{equation*}
$$

It can also be seen that, the polygon $B F D T$ is a rectangle, which enables us to calculate right angle side $B F$ as below,

$$
\begin{equation*}
B F=C D=C O \times \sin \Omega=B O \times \cos \alpha \times \sin \Omega=\cos \alpha \sin \Omega \tag{11}
\end{equation*}
$$

From equation (10) and (11), the azimuth angle from due south at point $A$ can be calculated as below,

$$
\begin{equation*}
\sin (\measuredangle B G F)=\frac{B F}{B G}=\frac{\cos \alpha \sin \Omega}{\cos \Psi} \tag{12}
\end{equation*}
$$

Similarly, in order to conveniently use $\omega t$ to represent the hour angle, the following formula is considered,

$$
\begin{equation*}
\sin \Omega=\sin (\pi-\measuredangle Q O E)=\sin (\measuredangle Q O E)=\sin \omega t \tag{13}
\end{equation*}
$$

Substitute equation (13) to (12), get,

$$
\begin{equation*}
\sin (\measuredangle B G F)=\frac{B F}{B G}=\frac{\cos \alpha \sin \omega t}{\cos \Psi} \tag{14}
\end{equation*}
$$

In our case, we redefine the azimuth angle to be from due east instead of due south, and denote the angle with symbol $\Phi$. It can be seen that $\Phi$ is the complementary angle of $\measuredangle B G F$, thus,

$$
\begin{equation*}
\sin (\measuredangle B G F)=\sin \left(\frac{\pi}{2}-\Phi\right)=\cos \Phi \tag{15}
\end{equation*}
$$

Substitute equation (15) into (14), get the final azimuth equation,

$$
\begin{equation*}
\cos \Phi=\frac{\cos \alpha \sin \omega t}{\cos \Psi} \tag{16}
\end{equation*}
$$

Where,

$$
\begin{array}{rlr}
\Phi & =\text { solar azimuth angle from due east } & (\mathrm{rad}) \\
\alpha & =\text { solar declination angle } & (\mathrm{rad}) \\
\omega & =\text { earth angular velocity } & \left(\frac{\pi}{12} h^{-1}\right) \\
t & =\text { local time }[0,24) & (h)  \tag{h}\\
\Psi & =\text { solar altitude angle } & (\mathrm{rad})
\end{array}
$$

