

AM111 Lectures**Sections****Office Hours**

- ☐ Week 5
- ☐ Week 6
- ☐ Week 7
- ☐ Week 8
- ☐ Week 9 Spring Break!
- ☐ Week 10
- ☐ Week 11

Apr. 10th

- ☐ *Apr. 11th*
 - ☐ Last time:
 - ☐ How to get better accuracy than the Euler method
 - ☐ Higher-order Taylor methods
 - ☐ Multistage single-step methods
 - ☐ How to derive them
 - ☐ Examples: 2nd order and 4th order Runge-Kutta
 - ☐ Error Control
 - ☐ If we know that the local error is $\propto h^{m+1}$ and the error is e for a particular h , then given a global error tolerance ϵ , we can change the step-size from $h \rightarrow qh$ s.t. the new error $|q^m e| \simeq \epsilon$

This implies

$$q \simeq \left| \frac{1}{2e} \right|^{1/m}$$

- ☐ One approach: Step doubling

over h : $e(h) = kh^{m+1}$

$$e(h/2) = 2k \left(\frac{1}{2} \right)^{m+1} h^{m+1}$$

$$y_{n+1}(h) - y_{n+1}(h/2) = \left[1 - \left(\frac{1}{2} \right)^m \right] kh^{m+1} = [2^m - 1] e(h/2)$$

- ☐ For a 4th order method:

$$y_{n+1}(h) - y_{n+1}(h/2) = 15e(h/2)$$

We estimate the error made as:

$$e(h/2) = \frac{y_{n+1}(h) - y_{n+1}(h/2)}{2^m - 1}$$

- ☐ This works nicely enough, and is easy to write, but has large overhead.
- ☐ A more efficient (modern) way: Embedded Runge-Kutta method
 - ☐ Use two methods of different order (remember ode45!)
 - ☐ (1) $\tilde{e} = y(t_{n+1}) - \tilde{y}_{n+1}$ this is $O(h^{m+2})$
 - ☐ (2) $e = y(t_{n+1}) - y_{n+1}$ is $O(h^{m+1})$
 - ☐ $e = \tilde{e} + \tilde{y}_{n+1} - y_{n+1} \simeq \tilde{y}_{n+1} - y_{n+1}$
 - ☐ The way we do this efficiently is to use the slopes needed to calculate the lower-order method to help calculate the higher-order method.
 - ☐ See section 7.5 in Moler's NCM. for an example of a 2nd order method with a 3rd order error estimate.
 - ☐ Implementation:
 - ☐ Calculate y_{n+1} using the lower order method, and $\tilde{y}_{n+1}$ using the higher order method. Then:

$$q = \left| \frac{1}{2(\tilde{y}_{n+1} - y_{n+1})} \right|^{1/m}$$
 - ☐ $h = qh$
 - ☐ if $q < 1$, repeat with the new h

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- ☐ else ($q \geq 1$), continue with new h
- ☐ Examples:
 - ☐ $F = @(t,y) 0$; ode23(F,[0 10],1)
 - ☐ $F = @(t,y) t$; ode23(F,[0 10],1)
 - ☐ $F = @(t,y) y$; ode23(F,[0 10],1)
 - ☐ $F = @(t,y) -y$; ode23(F,[0 10],1)
 - ☐ $F = @(t,y) \sin(t)$; ode23(F,[0 10],1)
 - ☐ Cruddy result --Boost tolerance!
 - ☐ $opts = odeset('RelTol', 1e-6)$;
 - ☐ $F = @(t,y) \sin(t)$; ode23(F,[0 10],1,opts)
 - ☐ Much better!
- ☐ Pendulum Equation: $\ddot{\theta} = -\frac{g}{L} \sin \theta$
- ☐ Watch the tolerance!
- ☐ Lorenz Equation:
 - ☐ $\dot{y}_1 = -\beta y_1 + y_2 y_3$
 - $\dot{y}_2 = -\sigma y_2 + \sigma y_3$
 - $\dot{y}_3 = -y_2 y_1 + \rho y_2 - y_3$
- ☐ Multistep Methods
 - ☐ Idea is to make use of previously computed y_k , $k \leq n$, to compute y_{n+1}
 - ☐
 - ☐ $\frac{dy}{dt} = f(t,y)$
 - ☐ $y_{n+1} = y_n + \int_{t_n}^{t_{n+1}} f(t,y) dt$
 - ☐ Idea: $y_{n+1} = y_n + \int_{t_n}^{t_{n+1}} P(t,y) dt$ We will choose different P's to approximate f .
 - ☐ First order scheme: $P_1(t) = \text{constant}$
 - ☐ $y_{n+1} = y_n + \Delta t f(t_n, y_n)$ Euler method!
 - ☐ 2nd order method:
 - ☐ $P_2(t) = f_{n-1} + \frac{f_n - f_{n-1}}{\Delta t} (t - t_n)$
 - ☐ Inserted into the integral equation, this yields:
 - ☐ $y_{n+1} = y_n + \frac{\Delta t}{2} [3f(t_n, y_n) - f(t_{n-1}, y_{n-1})]$
 - ☐ This is the 2nd order Adams-Bashforth method.
 - ☐ In general, methods which don't involve y_{n+1} (i.e. explicit methods) are called Adams-Bashforth methods.
 - ☐ When y_{n+1} is used to determine P , it's called an Adams-Moulton method
 - ☐ 1st order A-M:
 - ☐ $P_1(t) = \text{const} = f(t_{n+1}, y_{n+1})$
 - ☐ $y_{n+1} = y_n + \Delta t f(t_{n+1}, y_{n+1})$
 - ☐ This is the backward Euler Method!
 - ☐ 2nd order A-M:
 - ☐ $P_2(t) = f_n + \frac{f_{n+1} - f_n}{\Delta t} (t - t_n)$
 - (note similarity to the 2nd order A-B method.)
 - ☐ $y_{n+1} = y_n + \frac{\Delta t}{2} [f(t_{n+1}, y_{n+1}) + f(t_n, y_n)]$
 - ☐ This is the Trapezoidal Method.
 - ☐ Predictor-Corrector Method
 - ☐ 2nd order example

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- ☐ Predictor (A-B) :
$$y_{n+1}^p = y_n + \frac{\Delta t}{2} [3f_n - f_{n-1}]$$
- ☐ Corrector (A-M):
$$y_{n+1} = y_n + \frac{\Delta t}{2} [f(t_{n+1}, y_{n+1}^p) + f(t_n, y_n)]$$
- ☐ Used to model the orbital mechanics of spacecraft.
- ☐ Can we achieve arbitrary accuracy?
 - ☐ Of course not.
 - ☐ Suppose we integrate over an interval of length $L = t_f - t_0$
 - ☐ The # of steps we take to do this integral is $N=L/h$.
 - ☐ The roundoff error for each step is ϵ .
 - ☐ So the total roundoff error is $< N\epsilon$ (more realistically, this scales as $\sqrt{N\epsilon}$)
 - ☐ The total error is thus: $Ch^m + \frac{L}{h}\epsilon$ at mth order. If we make h too small, then this blows up.
- ☐ Stiffness of Problems:
 - ☐ What are some examples of stiff ODEs?
 - ☐ $\frac{dy}{dt} = \Lambda y$
 - ☐ $y(t=0) = \begin{bmatrix} 1 \\ 1 \end{bmatrix}$
 - ☐ $\Lambda = \begin{bmatrix} -100 & 1 \\ 0 & -\frac{1}{10} \end{bmatrix}$
 - ☐ Using the forward Euler method yields a solution which blows up at large time!
 - ☐ If we try the backward Euler method, we find that everything's just spiffy. More next time.
- ☐ Apr. 13th
- ☐ Round-off Errors
 - ☐ Over the interval $[t_0, t_f]$ of length $L = t_f - t_0$, the discretization error is Ch^p , and the roundoff error is $\frac{L}{h}\epsilon$.
 - ☐ For various orders, (if $L = 20$, $C = 100$, $\epsilon = 2^{-52}$)
 - ☐ $p=1$: roundoff error becomes important at $N = 4.5 \times 10^{17}$
 - ☐ $p=3$: " $N=5,647,721$
 - ☐ $p=5$: " $N=37,285$
 - ☐ $p = 10$: " $N=864$
- ☐ Stiff problems
 - ☐ example: $\frac{dy}{dt} = \Lambda y$
 - ☐ $y(0) = y_0$
 - ☐ $\Lambda = \begin{bmatrix} -100 & 0 \\ 0 & -\frac{1}{10} \end{bmatrix}$. This problem is unstable with the forward Euler method.
 - ☐ has two eigenvalues
 - ☐ $\lambda_1 = -100$
 - ☐ $\lambda_2 = -1/10$
 - ☐ The forward Euler method is only stable inside the unit circle centered on $z=-1$ on the complex plane.
 - ☐ So if we look at the stability condition:
 - ☐ $|1 + \lambda \Delta t| < 1$

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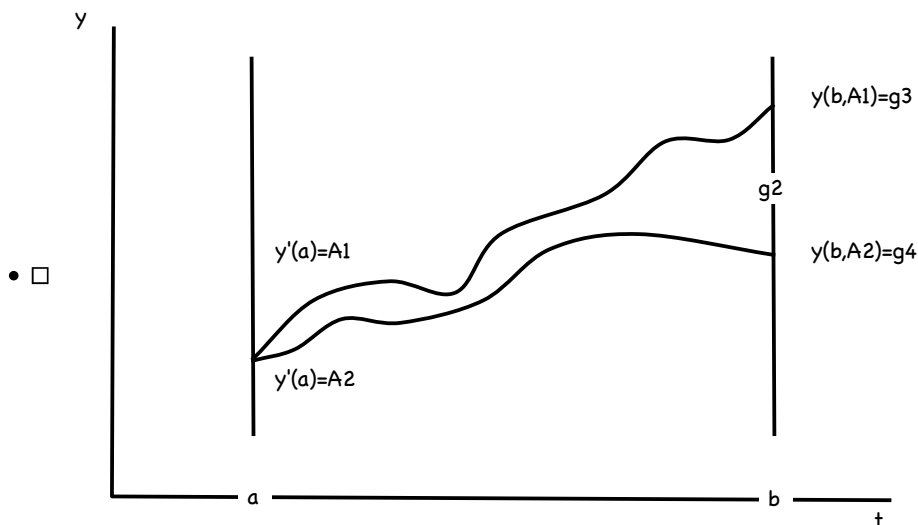
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- $\square \quad |1 + \lambda_1 \Delta t| = 9 < 1$, if our timestep is too large. (as it was when we tried this before)
- $\square$ If we make a smaller timestep, so that
- $\square \quad |1 + \lambda_1 \Delta t| = 1$ (when $\Delta t = 0.02$), we find that the solution using the forward Euler method doesn't blow up!
- $\square$ Even if the timestep is slightly too large ($\Delta t = 0.021$) the problem will eventually blow up. (although it does so more slowly than before)
- $\square$ **Definition:** A problem is "stiff" if the solution being sought varies slowly, but there are nearby solutions that vary rapidly, so that the numerical methods must take small steps obtain satisfactory results.
 - $\square$ **Alternative Definition:** A problem is stiff if its numerical solution by some methods requires (perhaps in only a portion of the solution interval) a significant depression of the step-size to avoid instability.
- $\square$ The **stiffness ratio** is the ratio of the largest and smallest (in modulus) eigenvalues of a linear system (for a general problem, these are the eigenvalues of the Jacobian matrix)
 - $\square$ The stiffness ratio of our example problem was 1000.
 - $\square$ For comparison, consider the world record stiffness of 10^{31} (from cosmological Big-Bang simulation)
- $\square$ Looking back at section 7.12 in the NCM, we now understand ode45, ode23, and ode113. What about ode15s? It uses backward differentiation formulas (BDFs)
 - $\square$ BDF
 - $\square \quad \frac{dy}{dt} = f(t, y)$
 - $\square \quad y_{n+1} = y_n + \int_{t_n}^{t_{n+1}} f(t, y) dt \simeq y_n + \int_{t_n}^{t_{n+1}} P(t, y) dt$
- $\square$ ode23s is based on a modified Rosenbrock formula of order 2
 - $\square$ Rosenbrock formula:
 - $\square$ This is a generalized implicit Runge-Kutta formula
 - $\square \quad s_i = f(t, \sum_{k=1}^{i-1} \beta_{i,k} s_k)$ is an explicit method. If we let the sum go over all the slopes,
 - then
 - $\square \quad s_i = f(t, \sum_{k=1}^i \beta_{i,k} s_k)$ we have an implicit method.
- $\square$ Consider the problem of the size of an expanding flame. The flame depends upon oxygen to exist, so it can grow at a rate proportional to the rate at which oxygen is available to it. If the ball is of radius y , then the rate of injection of oxygen is proportional to y^2 . Oxygen consumption should go as the volume of the flame, or y^3 . Thus
- $\square \quad \dot{y} = y^2 - y^3$
- $\square$ let $y(0) = \delta$, and then run the simulation from $0 \leq t \leq 2/\delta$
- $\square$ We notice that this problem is solved much faster with implicit methods than by explicit methods. (recall section, and NCM)
- $\square$ Why is this so? To find out, let's linearize the equations (find the Jacobian)
 - $\square \quad J = f_y = 2y - 3y^2$ Near equilibrium, $y=1$.
 - $\square \quad \dot{y} = J|(y - y_c) = -1(y - 1)$
 - $\square \quad \lambda = -1$, so the timestep must be $\Delta t \leq 2$, which seems big, but the span we're trying to cover is huge, so this severely constrains the span we can solve this over.
 - $\square$ Implicit methods, by contrast, allow us to take arbitrarily huge timesteps once we're near equilibrium. They're much better for problems like these.
- $\square$ Summary:
 - $\square$ Single-step (multistage) methods
 - $\square$ Multistep methods
 - $\square$ useful for smooth problems that require high accuracy and where evaluations of $f(t, y)$ are expensive
 - $\square$ Explicit methods are easier to implement

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- ☐ Implicit methods have larger regions of stability
- ☐ Boundary Value Problems:
 - ☐ A 2nd order BVP is (in general)
 - ☐ $\frac{d^2y}{dt^2} = f(t, y, \frac{dy}{dt})$
 - ☐ on $t \in [a, b]$ with the general boundary conditions
 - ☐ $\alpha_1 y(a) + \beta_1 \frac{dy(a)}{dt} = \gamma_1$
 - ☐ $\alpha_2 y(b) + \beta_2 \frac{dy(b)}{dt} = \gamma_2$
 - ☐ We don't have enough information to set up an initial condition to run forward from, so what do we do?
 - ☐ Shooting Methods:
 - ☐ guess that $y(a) = \gamma_1, y'(a) = A_1$. We then run the initial condition problem forward and find out that this leads to $y(b, A_1) = \gamma_3$. This is a tad off (perhaps bigger than γ_2), so we try another guess, $y(a) = \gamma_1, y'(a) = A_2$. This gives us another result $y(b, A_2) = \gamma_4$, which might be smaller than γ_2 . The general problem is to find an argument A s.t.



- ☐ $y_b(A) - \gamma_2 = 0$
 - ☐ This is a zero-finding problem!
- ☐ Finite Difference Methods
 - ☐ For example, the Schrödinger equation
 - ☐ $\frac{d^2y}{dt^2} + [V(t) - E]y = 0$
 - ☐ $y(0) = 0$
 - ☐ $y(1) = 0$
 - ☐ $\frac{d^2y}{dt^2} = \frac{y(t + \Delta t) - 2y(t) + y(t - \Delta t)}{\Delta t^2}$
 - ☐ so that
 - ☐ $\frac{y_{n+1} - 2y_n + y_{n-1}}{\Delta t^2} + (V_0 - E)y_n = 0$, with $y_1 = y_N = 0$.
 - ☐ This can be written in matrix form as

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$$\bullet \square \begin{bmatrix} -2 + V_0 \Delta t^2 & 1 & 0 & 0 & \dots & 0 \\ 1 & -2 + V_0 \Delta t^2 & 1 & 0 & \dots & 0 \\ 0 & \ddots & \ddots & \ddots & 0 & 0 \\ 0 & 0 & \dots & 0 & 1 & -2 + V_0 \Delta t^2 \end{bmatrix} \begin{bmatrix} y_2 \\ y_3 \\ \vdots \\ y_{N-1} \end{bmatrix} = E \begin{bmatrix} y_2 \\ y_3 \\ \vdots \\ y_{N-1} \end{bmatrix}$$

- ☐ Let's try an example

- ☐ $\frac{d^2 y}{dt^2} = y^2 - 1$

- ☐ $y(0) = 0, y(1) = 1$

- ☐ Using a shooting method

- ☐ $y(0) = 0, y'(0) = 1 \leftarrow$ a guess

- ☐ This leads to $y(t=1) - 1 \simeq .45$

- ☐ Trying $y(0) = 0, y'(0) = 2$

- ☐ This leads to $y(t=1) - 1 \simeq$ something less than 0

- ☐ Thus we've bracketed the zero, and we can just solve this problem using whatever our favorite root finding method is.

- ☐ We can also try a finite difference method:

- ☐ This forms a linear system $A(y)y = b$ (the coefficient matrix A depends upon y)

- ☐ This requires some iteration to solve.

- ☐ $\frac{y_{n+1} - 2y_n + y_{n-1}}{\Delta t^2} = y_n^2 - 1$, which, in matrix form is

$$\bullet \square \begin{bmatrix} -2 & 1 & 0 & 0 & 0 & \dots \\ 1 & -2 & 1 & 0 & 0 & 0 \\ 0 & 1 & -2 & 1 & 0 & 0 \\ \ddots & \ddots & \ddots & \ddots & \ddots & \vdots \\ \dots & 0 & 0 & 0 & 1 & -2 \end{bmatrix} \begin{bmatrix} y_2 \\ y_3 \\ y_4 \\ \vdots \\ y_{N-1} \end{bmatrix} + \begin{bmatrix} 0 \\ 0 \\ 0 \\ 0 \\ 0 \\ 1 \end{bmatrix} = A y + b$$

- ☐ We have $A y + b = \Delta t^2 c(y)$

- ☐ This can be solved by linear iteration.

- ☐ Newton's method:

- ☐ Residue: $A y + b - (\Delta t)^2 (y^2 - 1)$

- ☐ Jacobian: $A - 2(\Delta t)^2 y$

- ☐ $J \Delta y = -\text{Residue}$

- ☐ Alternatively, we could just use the MATLAB function **bvp4c**

- ☐ To use, we first:

- ☐ (a) write the ODE in standard form

- ☐ (b) write a function to describe your boundary condition

- ☐ $y_2 = \dot{y}_1$

- ☐ $\dot{y}_1 = y_2, \dot{y}_2 = y_1^2 - 1$

- ☐ $y_1(0) = 0$

- ☐ $y_1(1) - 1 = 0$

- ☐

Apr. 14th

- ☐ Week 12

- ☐ Week 13

- ☐ Week 14